

# The Formation of Long Carbon Chain Molecules during Laser Vaporization of Graphite

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**Abstract:** Graphite is laser vaporized into a He carrier gas containing various simple molecules such as H<sub>2</sub>, H<sub>2</sub>O, NH<sub>3</sub>, and CH<sub>3</sub>CN, supersonically expanded, and skimmed into a molecular beam, and the beam is interrogated by photoionization time-of-flight mass spectrometry. Without added reactants in the He carrier gas, C<sub>n</sub> species up to n = 130 are readily observed. Two distributions separated at about n = 40 appear to be present with the low n species the focus of this work. In the presence of added reagents, new species appear as a result of reaction. These are satisfactorily explained on the basis that a significant proportion of the C<sub>n</sub> species initially formed are reactive radicals with linear carbon chain structures which can readily add H, N, or CN at the ends to form relatively stable polyynes or cyanopolynes. Some of the cyanopolynes detected have also been observed in the interstellar medium, and circumstellar carbon condensation processes in the atmospheres of carbon-rich stars similar to those studied here are suggested as possible synthetic sources.

At present the most successful schemes for interstellar molecule formation are those which propose that bimolecular ion-molecule processes occur in the cold gaseous dark clouds where they are detected.<sup>1,2</sup> Although the abundances of many small molecules can be satisfactorily explained by this theory,<sup>3,4</sup> it is far from clear that they offer a plausible route to the long-chain polyynes, HC<sub>n</sub>N (n = 3-11), that have been observed in significant amounts.<sup>5-9</sup> Even some of the largest and most complete attempts to explain the range of molecules observed are not able to extend much past HC<sub>3</sub>N, i.e., species such as HC<sub>5</sub>N cannot be included in a totally satisfactory manner without a great increase in the scale of the calculation.<sup>10</sup> Clearly other mechanisms for the formation of these molecules need to be explored. One plausible mechanism might be the formation of these species by bombardment of carbonaceous grains by energetic neutral atoms. Devienne and Teisseire<sup>11</sup> have carried out an interesting experimental study which produced several of the long carbon chain species of interest here by such bombardment of graphite.

The long-chain molecules are detected not only in cold molecular clouds but also in the atmospheres of certain stars such as the cool carbon rich red giant IRC+10216.<sup>9,12</sup> In addition to ejecting the polyynes (and many other molecules) into the general interstellar medium, such stars also appear to eject carbon particles.<sup>12,13</sup> It thus seems appropriate to consider an alternative scenario for the formation of carbon chain molecules which involves conditions that relate more closely to those that occur in stellar atmospheres in which carbon is condensing.<sup>12</sup>

The present study was initiated primarily to explore the particular possibility that long carbon chain molecules, such as the cyanopolynes that have actually been detected in the interstellar medium (ISM), can form under conditions in which carbon vapor nucleates in the presence of hydrogen and nitrogen, conditions similar to those which are expected to occur in the atmosphere of a carbon star. With this objective a series of experiments was undertaken in which carbon was laser vaporized into a helium stream containing various added gases, allowed to cluster and react, and studied by photoionization mass spectrometry. The astrophysical implications of the present experiments will be considered in more detail elsewhere<sup>14</sup> and will be described here only briefly.

There have been a number of important previous studies of carbon clustering. Some early mass spectrometric measurements<sup>15,16</sup> detected carbon clusters with up to 33 atoms in carbon vapor. More recent experiments have shown that much larger clusters with more than 100 atoms can also form,<sup>17-19</sup> and convincing evidence for linearity of the even clusters, C<sub>m</sub>, in the size range 1 < n < 24 has been obtained.<sup>17</sup> Study of the larger clusters

has shown that some of the larger clusters, particularly C<sub>60</sub>, have some unique properties<sup>19-22</sup> which are most readily explained if they are inert hollow spheroidal shell molecules. Moreover, a laser vaporization cluster beam<sup>23</sup> study indicates that macroscopic particles also form, in addition to the long-chain molecules and large spheroidal clusters. These discoveries also give insight into the homogeneous nucleation mechanism governing the formation of carbonaceous particles such as soot and carbon black and the explanation which they provide for the resultant morphology of these particles.<sup>21</sup>

The work presently described concentrates on the smaller carbon clusters with from 6 to 30 atoms and some reactions which they undergo. There has been much conjecture about the structures of these clusters: whether or not, for instance, they are linear carbon chains or monocyclic rings.<sup>16</sup> Odd/even variations of intensity in mass spectrometric data have been interpreted in terms of linearity.<sup>16,24</sup> However, some of the mass spectrometric data

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indicated that clusters which had  $4n + 7$  ( $n = 1-4$ ) atoms might have special stability. This so-called " $\Delta n = 4$  effect" has been considered as possible evidence for ring structures.<sup>16,25</sup> However, there is convincing evidence that the even clusters with  $n < 24$  are linear<sup>17</sup> and our results strongly support this conclusion for at least a significant fraction of the species in this range detected.

### Experimental Section

A detailed description of the apparatus used in this study has been given previously.<sup>23,26-29</sup> A 2.5 cm diameter disk, cut from 0.3 mm sheet of high-density graphite, was mounted so that the vaporized products formed at its surface by a focused pulsed laser can be entrained in a pulse of He carrier gas just before it expands through the throat of a supersonic nozzle. The disk was rotated and translated during the experiments so that fresh surface was continuously exposed to the laser (30-40 mJ of 2nd harmonic Nd:YAG laser radiation focused to a 0.1-mm spot) which was fired to coincide with the peak in the He gas density over the target area. The gaseous products flowed via a channel (2 mm wide expanding after 13 mm to 4 mm for a further 10 mm) into a large vacuum chamber. The supersonic jet of He entrained clusters so produced was skimmed 30 cm downstream to form a beam which passed through two further stages of differential pumping into the ionization chamber of a time-of-flight (TOF) mass spectrometer. At this point the cluster population was interrogated by a second pulse of ionizing ArF excimer laser radiation (10 mJ at 6.4 eV), which could be fired to synchronize with the characteristic arrival times of the clusters, producing ions whose masses were determined by TOF mass spectrometry. The main timing sequence is controlled by an IBM-AT microcomputer interfaced to a microsecond delay generator via a CAMAC crate. In general an individual experiment yielded a mass spectrum averaged over a sequence of ca. 1000 cycles repeated at 0.11-s intervals.

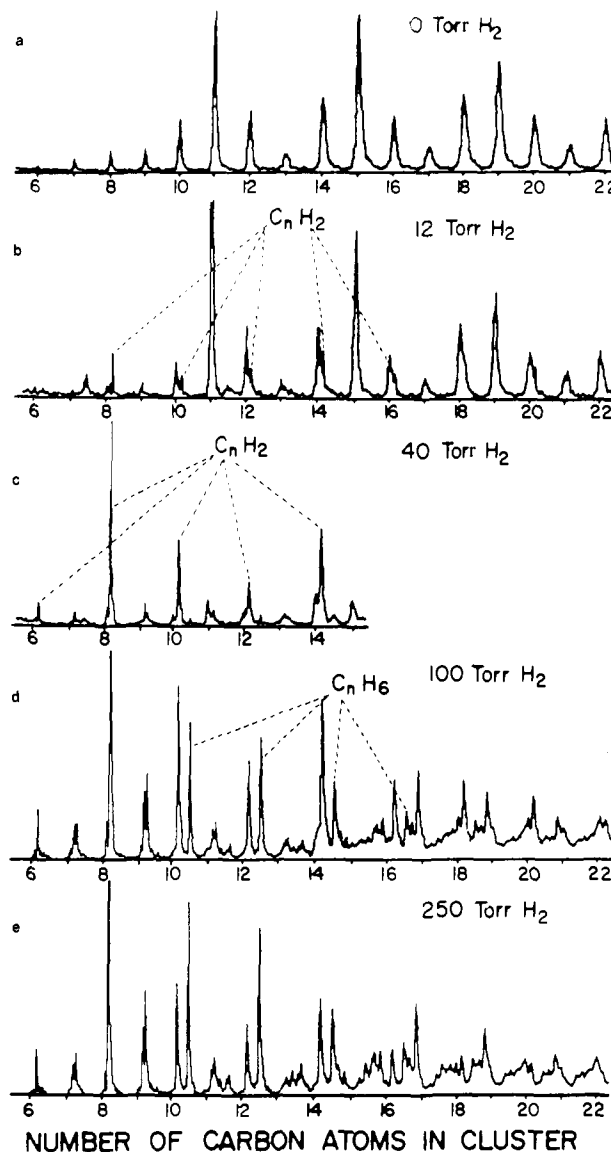
In the initial experiments pure He entraining gas was used, producing the basic carbon cluster distribution. To carry out the reaction studies, carefully controlled amounts of reactants were introduced into the He. In order to do this, the various gas flow rates were monitored by separate floating ball flowmeters and, in especially critical cases, a Hastings Teledyne digital flowmeter was also used to monitor the reactant. The  $H_2$  and  $N_2$  reactants as well as the He were filtered whereas the  $O_2/He$  mixture was used directly. Experiments with water were carried out by passing the He through a steel capsule containing cotton wool soaked in either  $H_2O$  or  $D_2O$ . The He backing pressure and capsule temperature were adjusted to achieve estimated water/He mixtures of ca. 1/1000.

### Results

A typical cluster mass distribution without added reactants (obtained using ArF laser photoionization) is depicted in Figure 1a. It should be noted that, as discussed above, clusters with more than 30 atoms are also observed, which behave in a different way<sup>19-22</sup> from the smaller ones which are considered in this study. The curious prominence of  $C_{11}$ ,  $C_{15}$ ,  $C_{19}$ , and, to some extent,  $C_{23}$  (the  $\Delta n = 4$  effect) is clearly evident. The ring hypothesis is considered further in the Discussion section.

In the following sections we describe the results obtained when various reagent gases are added to the gas stream. These gases are exposed to the very high plasma temperatures and are presumably partially decomposed into very reactive atomic and free radical species which are in turn the major reactants with the carbon clusters.

**Reactions with  $H_2$ .** The results of the reactions which occur when various amounts of  $H_2$  are entrained in the He pulse are presented in Figure 1, spectra b-e, in which the ratio of  $H_2:He$  is varied. When a small amount (12 Torr) of  $H_2$  is entrained (Figure 1b), relatively marked changes are detected immediately. The first noticeable result is that peaks appear at  $C_nH_2$  with  $n = 8, 10, 12,$  and  $14$ , and by the time 40 Torr have been introduced (Figure 1c), the dominance of  $C_{11}$  and  $C_{15}$  has been wiped out



**Figure 1.** (a) Mass spectrum of bare carbon clusters, from  $C_5$  to  $C_{22}$ , produced in the presence of a pure He pulse. (b) Result obtained when ca. 12 Torr of  $H_2$  is entrained in the He (ca. atmospheric pressure in general) pulse. The even  $n$  peaks grow +2 mass unit satellites producing peaks corresponding to  $C_8H_2$ ,  $C_{10}H_2$ ,  $C_{12}H_2$ ,  $C_{14}H_2$ ,  $C_{16}H_2$ , and  $C_{18}H_2$  which can be clearly identified. Note that  $C_8H_2$  is strong almost immediately even though  $C_8$  is very weak in spectrum a. (c) When ca. 40 Torr of  $H_2$  is entrained the mass spectrum contains essentially only hydrogenated products. Note that the odd carbon clusters which are still bare in spectrum b have now reacted, indicating that although they are reactive they are slightly less so than the even ones. The most striking changes are the following: the strength of the  $C_8H_2$  peak in view of the insignificance of the bare  $C_8$  cluster in spectrum a, the appearance of extra peaks for other even clusters, and the dramatic collapse of the  $C_{11}$  and  $C_{15}$  peaks. (d) In this scan 100 Torr of  $H_2$  were entrained in the He. We note that for the even clusters the  $C_nH_2$  peak for a given even  $n$  value gradually loses its dominance as  $n$  increases though the peak for  $C_{18}H_2$  is clearly identifiable. For  $n = 10$  and  $12$ ,  $C_nH_6$  are becoming the dominant peaks and for  $n = 16, 18,$  and  $20$  the  $C_nH_{10}$  peaks are becoming dominant. The odd  $n$  cluster peaks have been replaced by groups of weak peaks corresponding to a varied range of hydrogenated products. The range is narrow for low  $n$  where only a few sharp peaks are observed and wide at high  $n$  where unresolved broad signals are detected. (e) When 250 Torr of  $H_2$  are entrained, the dominant peaks (for even  $n$  species) are the following:  $C_6H_2$ ,  $C_8H_2$ ,  $C_{10}H_6$ , and  $C_{14}H_2$  and  $C_{14}H_6$ ,  $C_{16}H_{10}$ ,  $C_{18}H_{10}$ , and  $C_{20}H_{10}$ . In general no single species is dominant for the odd  $n$  set though  $C_7H_2$ ,  $C_9H_2$ , and  $C_9H_3$  do stand out.

and replaced by the dominance of the  $C_nH_2$  species ( $n$  even). In fact clear evidence for  $C_6H_2$  and  $C_8H_2$  is obtained whereas the bare  $C_6$  cluster is barely detectable and  $C_8$  is very weak indeed

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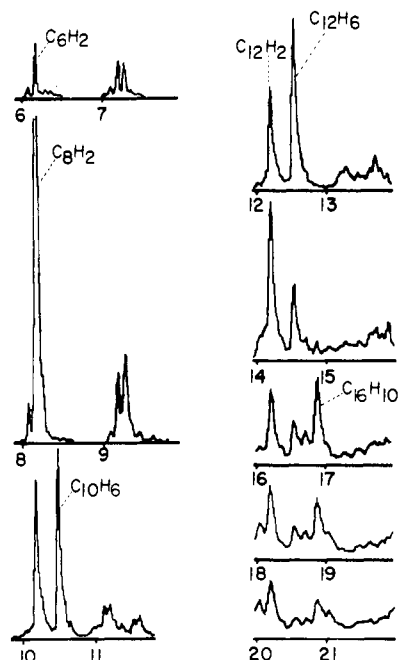
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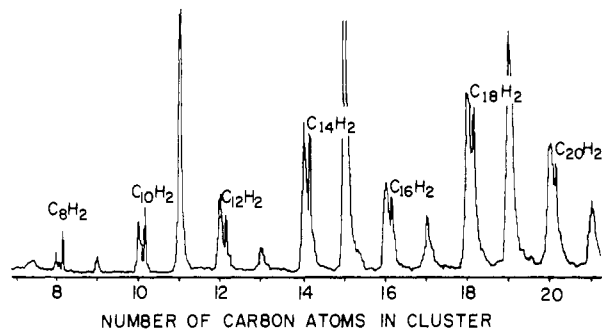
**Figure 2.** In this detailed re-representation of Figure 1, spectrum d, the products of hydrogenation can be compared as a function of increasing mass number. It is particularly striking that the first mass peak to be observed,  $C_nH_2$ , remains strong for all even  $n$  values.  $C_nH_6$  peaks are not observed for  $n < 10$ , and as  $n$  increases, this peak becomes weaker. Peaks at  $C_nH_8$  and  $C_nH_{10}$  appear at  $n = 14$  and  $20$ , the latter tending to share dominance with the  $C_nH_2$  features. The odd clusters behave in a much more irregular manner forming a range of hydrogenated products showing no particularly striking specificity.

in Figure 1a. In the 100-Torr  $H_2$  experiment, Figure 1d, strong peaks are also observed for  $C_{10}H_6$ ,  $C_{12}H_6$ , and  $C_{14}H_6$  (weaker features are also detected as discussed below). Even more striking changes take place in the case of the odd  $n$  clusters. In Figure 1c we see that the bare cluster peaks (including those for the dominant  $C_{11}$  and  $C_{15}$  species) have disappeared completely to be replaced by groups of relatively weak features due to sets of hydrogenated products.

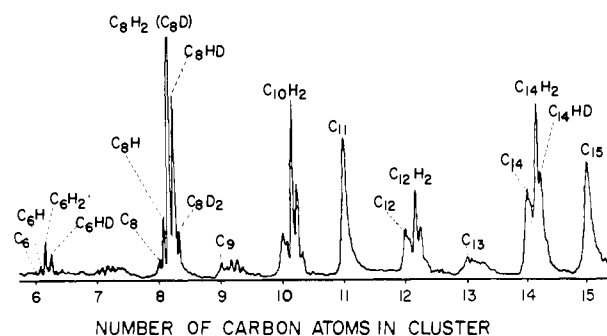
As the  $H_2$  concentration is increased (Figure 1, d and e),  $C_nH_6$  products begin to become prominent for  $n$  even and  $> 8$ . The essentially complete inertness of the  $C_8H_2$  species to further addition of hydrogen under these conditions is very noticeable. Some detailed aspects of the reaction are effectively highlighted in Figure 2 where the results of Figure 1d are displayed so that the products for successive odd and even  $n$  species can be compared in a vertical scheme. In this diagram, we find that, for even  $n$ , the  $C_nH_2$  signal is in general always relatively strong and that the  $C_nH_6$  signal, which only appears when  $n = 10$  (for which it is strong) or more, gradually diminishes as peaks at  $C_nH_8$  and particularly at  $C_nH_{10}$  appear as  $n$  increases from 14 to 20. The strength of the  $C_nH_2$  peak diminishes much more rapidly with increasing  $n$  at high  $H_2$  concentrations as can be seen from Figure 1, d and e. The dominant peaks for a given even  $n$  value are, in addition to  $C_nH_2$ , the following:  $C_{10}H_6$ ,  $C_{12}H_6$ ,  $C_{14}H_6$ ,  $C_{16}H_{10}$ ,  $C_{18}H_{10}$  and  $C_{20}H_{10}$ .

The behavior of the odd  $n$  species can also be followed as a function of increasing  $n$  in Figure 2 where it is seen to be rather more complicated. The general result is that these species react rapidly (though not quite as rapidly as the even clusters) to form a whole range of feasible hydrogenated hydrocarbons, including ones with odd numbers of hydrogen atoms.

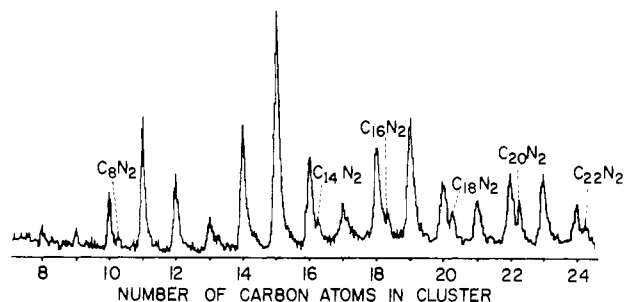
**Reactions with  $H_2O$ .** The results of the reactions that take place with trace quantities of water are nicely displayed in Figure 3 where unequivocal evidence for the formation of species with the formula  $C_nH_2$  for  $n = 10, 12, 14, 16, 18, 20$  is presented. The mass spectrum is very similar to that obtained with small amounts of  $H_2$  as depicted in Figure 1b. These two sets of data are clear evidence that the even  $n$  clusters are linear as the most obvious interpretation is that the products are polyyenes of structure



**Figure 3.** When water vapor is entrained in the He, very clear +2 mass unit satellites are observed on the even  $n$  peaks. These peaks are sharp relative to the bare cluster signals which may signify stability to ion fragmentation. In this experiment there is no obvious loss of the general underlying pattern of intensities which characterizes the bare cluster scan, Figure 1a. In this respect the water reaction appears to be very similar to the  $H_2$  experiment whose results are shown in Figure 1b. In particular we note that  $C_{11}$ ,  $C_{15}$ , and  $C_{19}$  remain dominant and apparently are fairly unreactive. In addition it appears that under these conditions the reaction has not proceeded further than the formation of  $C_nH_2$  species.



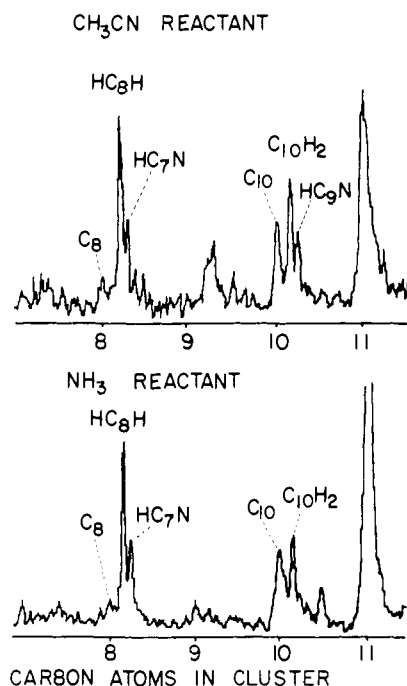
**Figure 4.** In this detailed scan the results obtained by using a  $D_2O/H_2O$  mixture entrained in the He pulse are presented. The degree of reaction is somewhat greater in this scan than in Figure 3. The even  $n$  peaks form the predicted species:  $C_n$ ,  $C_nH$ ,  $C_nD$ ,  $C_nH_2$ ,  $C_nHD$ , and  $C_nD_2$ . Although  $C_7$  and  $C_9$  have reacted and show satellites,  $C_{11}$  and  $C_{13}$  have not and  $C_{13}$  is, as usual, a weak broad feature.



**Figure 5.** In this scan the results of introducing  $N_2$  into the He pulse are shown. Although the general pattern characteristic of the bare carbon cluster distribution is detected, one observes that the even species peaks are accompanied by satellites at +4 mass units. These correspond to the formation of the symmetric dicyanopolyyenes,  $N\equiv C-(C\equiv C)_n-C\equiv N$ .

$H(C\equiv C)_mH$ . Further confirmation of this interpretation is furnished by the results of the mixed  $D_2O/H_2O$  experiment shown in Figure 4. Here the peaks are readily assigned as corresponding to  $C_n$ ,  $C_nH$ ,  $C_nH_2 + C_nD$ ,  $C_nHD$ , and  $C_nD_2$  as expected.

**Reactions with  $N_2$ .** The results of the reaction with  $N_2$  are shown in Figure 5 where each even  $n$  carbon peak is seen to develop a +4 mass unit satellite. This observation is readily interpreted as due to the formation of the dicyanopolyyenes, presumably produced by N atom addition to the two ends of the even carbon atom chains. The data are thus convincing evidence for the production of  $N\equiv C-(C\equiv C)_m-C\equiv N$  molecules, with  $m = 7-10$ , and also a further excellent indication that linear carbon chains are present.



**Figure 6.** In the upper scan is depicted the result of introducing  $\text{CH}_3\text{CN}$  into the He pulse. This molecule is expected to provide both H atoms and CN radicals which are expected to react with the even carbon chains to produce the polar cyanopolyynes. In this scan the  $\text{C}_8$  and  $\text{C}_{10}$  peaks are accompanied by peaks at +2 and +3 mass units consistent with polyyne formation and also +3 mass units consistent with cyanopolyne formation. The latter peaks belong to  $\text{HC}_7\text{N}$  and  $\text{HC}_9\text{N}$ , both of which have been detected in the interstellar medium. In the lower trace the results of introducing  $\text{NH}_3$  into the He are depicted. In this case H and N atoms are expected to react with the odd carbon cluster population to produce the cyanopolyynes. Clear evidence for  $\text{HC}_7\text{N}$  is observed as well as for  $\text{HC}_8\text{H}$  and  $\text{HC}_{10}\text{H}$  although the  $\text{HC}_9\text{N}$  signal is very weak.

**Reactions with  $\text{CH}_3\text{CN}$  and  $\text{NH}_3$ .** The reactions between the carbon clusters and reactants that can introduce hydrogen and nitrogen into the system at the same time offer the possibility of producing the polar cyanopolyynes such as  $\text{HC}_7\text{N}$  and  $\text{HC}_9\text{N}$  that actually have been detected in the ISM. In addition the interesting possibility of producing them from either the odd or the even  $n$  bare carbon cluster populations presents itself. In the upper trace in Figure 6 the result of introducing  $\text{CH}_3\text{CN}$  into the He is depicted which shows clear evidence for the production of  $\text{HC}_8\text{H}$ ,  $\text{HC}_7\text{N}$ ,  $\text{HC}_{10}\text{H}$ , and  $\text{HC}_9\text{N}$ . In the plasma it is likely that CN radicals will be formed in significant quantities as well as H atoms. Thus it is likely that a significant fraction of polar cyanopolyynes is produced by addition of H and CN to opposite ends of an even linear cluster population. In the lower trace of Figure 6 is depicted the result of adding  $\text{NH}_3$  which shows similar products, though the  $\text{HC}_9\text{N}$  peak is rather weak. In this case it is possible that the cyanopolyynes are produced by addition of H and N atoms to opposite ends of odd numbered carbon chain species (if such exist).

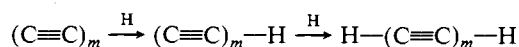
### Discussion

In this study, there are two major sources of concern in interpreting the observations. First, we are assuming that the dominant reactions involving entrained reagents take place after carbon has clustered so that the detected products are reliable probes of the properties of the pure carbon clusters. However, the possibility exists that H atom reactions take place at an early stage in the clustering process and produce intermediates which guide the clustering process toward the formation of linear polyynes. This seems unlikely because the carbon plasma at the point of ejection is constricted at the high He pressure (ca. 1 atm), diffusion is slow, and the carbon clustering reactions are very rapid in the resulting plume of concentrated carbon vapor.

Second, in the interpretation of some of the pertinent mass spectrometric observations we have assumed little fragmentation. As is discussed below, our results suggest that fragmentation may

be extensive for the bare clusters and thus may be responsible for the  $\Delta n = 4$  effect. The mass spectra of the even  $n$  reaction products, however, seem to consist largely of parent ions, since the major product ion peaks can be accounted for by saturated valence compounds. In summary, we believe that the peaks for the even  $n$  products are reliable probes of the original neutral bare parent clusters.

The major specific result of this series of experiments is highlighted by the result of the reaction with water depicted in Figure 3 where unequivocal evidence that there are only two highly reactive monovalent sites on the even bare clusters from  $n = 8$ –20 is presented. This result is most readily explained if the initial products formed from the even  $n$  clusters are the stable polyynes resulting from the addition of a hydrogen atom to each end of a polyyne chain diradical according to the scheme

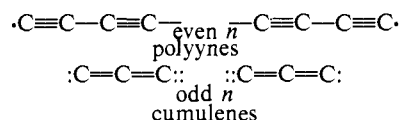


The results of the reactions with small amounts of hydrogen, Figure 1b, are much the same, suggesting further confirmation of this proposed explanation. It is interesting to note that the  $\text{C}_n\text{H}_2$  peaks are quite sharp relative to the bare clusters in Figure 3. One possible reason for this is that the bare cluster ions are metastable and fragment in the flight tube whereas the polyynes, which are relatively stable gas-phase species, produce parent ions which are not fragmenting.

At higher hydrogen concentrations, additional hydrogen atoms add to the larger even clusters four at a time. This can be rationalized in terms of linear polyyne chains in which triple bonds are being reduced to a single bond. This may be a consequence of a tendency of triple bonds to exhibit somewhat lower reactivity than double bonds under these conditions.

A further interesting observation on the general reactivity behavior of the various clusters is the inertness of  $\text{C}_8\text{H}_2$  to further reaction. In Figure 1 we see that the bare  $\text{C}_8$  signal is weak in Figure 1a and by Figure 1b it is the  $\text{C}_8\text{H}_2$  peak which is the dominant feature. At the highest hydrogen pressure the whole spectrum is dominated by the  $\text{C}_8\text{H}_2$  signal as this species does not seem to undergo further reaction. This suggests that under these conditions, conditions under which the longer chains are hydrogenated,  $\text{C}_8\text{H}_2$  (and to some extent also  $\text{C}_6\text{H}_2$ ) is extremely stable. This behavior is intriguing as it is in apparent contrast with the behavior of the higher even  $n$   $\text{C}_n\text{H}_2$  homologues. As indicated in Figures 1 and 2, the clusters  $\text{C}_{12}$ ,  $\text{C}_{14}$ , and  $\text{C}_{16}$  clearly give rise to prominent  $\text{C}_n\text{H}_6$  peaks as well as  $\text{C}_{14}$ ,  $\text{C}_{16}$ ,  $\text{C}_{18}$ , and  $\text{C}_{20}$  higher products. This observation indicates that there is a change in behavior after  $n = 8$ ; however, the reason is still not certain. It may be possible that there is more than one isomer of  $\text{C}_n$ ,  $n > 9$  (as was found<sup>30</sup> recently for  $\text{C}_7^+$ ) with only a portion of the  $n = 10$  and higher clusters being linear.

Comparing the odd cluster reactivities with those of the even ones, the mass spectra for the reaction products appear to display an odd/even alternation. Although it is possible that the odd clusters are not linear, the expected chemical differences between even  $n$  polyynes and odd  $n$  cumulenes



can explain the differences between odd and even cluster behavior. Cumulenes would certainly be expected to yield a more complicated range of initial products. Pitzer and Clementi<sup>24</sup> suggest that the even chains also possess cumulene-type bonding in their ground states; however, the present results are consistent with polyyne-like reactivities.

The  $\Delta n = 4$  effect is strikingly depicted in Figure 1a where the characteristic pattern for the mass spectrum of the bare clusters is shown. On the basis of the hydrogen reaction results, however,

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we observe that the bare  $C_9$ ,  $C_{13}$ , and  $C_{17}$  peaks are replaced by hydrogenated product peaks which are similar both qualitatively and quantitatively to those which result from  $C_{11}$ ,  $C_{15}$ , and  $C_{19}$  even though the bare peaks of the former set are at least an order of magnitude weaker than those of the latter. In fact the resulting features vary rather gradually in pattern and intensity across the range of odd  $n$  products. Because we cannot interpret the hydrogenated peaks for odd  $n$ , it is not clear whether we are observing parent ions. However, the smooth variation of ion intensities with  $n$  (odd) suggests that the neutral precursors are more closely related in character than the original  $\Delta n = 4$  pattern implies. This suggests that the characteristic intensity pattern of the mass spectrum observed for the bare clusters in Figure 1a may be dominated by ion fragmentation and is not a true diagnostic of the neutral cluster relative abundance. The  $\Delta n = 4$  effect may be due in some way to ring structures—probably of the ions ultimately detected.

After the work detailed here was completed the interesting experiments of Devienne and Teisseire<sup>11</sup> were brought to our attention. Their observations and ours are complementary as they have shown that, given carbonaceous grain (graphitic) particles,

a range of interstellar species can be produced by grain bombardment with high-energy neutrals.

In the present experiments we conclude that such species can be formed more or less simultaneously with carbonaceous grains. Although the conditions in our experiments differ from those in stellar atmospheres, it is highly likely that the main features of the carbon nucleation process are similar. Therefore, we feel that these experiments demonstrate a simple way in which the long-chain cyanopolynes observed by radioastronomy can be formed in regions where carbon is condensing, and in the absence of any alternative similarly convincing interstellar process, the results suggest a different source of these species from those heretofore considered.<sup>12,11</sup>

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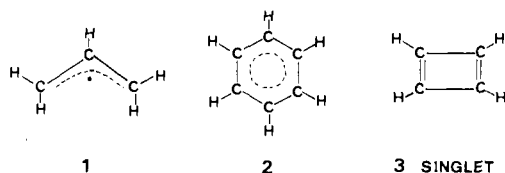
## Is Delocalization a Driving Force in Chemistry? Benzene, Allyl Radical, Cyclobutadiene, and Their Isoelectronic Species

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**Abstract:** The VB correlation diagram model (Figure 1) is used to answer the title question. It is shown that *only atoms that form weak two-electron bonds with low triplet excitation energies may generate delocalized species that are stable toward a localizing distortion.* Electronic delocalization is, then, seldom expected to be a significant driving force in chemistry. By this principle, the  $\pi$ -components of delocalized species, like  $C_6H_6$  and  $C_3H_5$ , are predicted to be distortive electronic systems that are trapped, within "rigidly" symmetric  $\sigma$ -frames, and are thereby delocalized despite their opposite inherent tendency. The predictions are examined by means of ab initio investigations at the levels of STO-3G, 6-31G, and 6-311G with extensive correlation (CI) calculations (up to  $6 \times 10^6$  determinants).  $\sigma$ - $\pi$  energy partitions show that the  $\pi$ -components of  $C_6H_6$  and  $C_3H_5$  are indeed distortive much like the  $\pi$ -electrons of  $C_4H_4$ , and all the  $\pi$ -components resemble, in turn, their isoelectronic  $H_n$  ( $n = 3, 4, 6$ ) species in the common reluctance to adopt geometries that lead to electronic delocalization. *Electronic delocalization in  $C_3H_5$  and  $C_6H_6$  turns out to be a byproduct of the  $\sigma$ -imposed geometric symmetry and not a driving force by itself.* The  $\pi$ -distortive propensities are shown to coexist harmoniously with the thermochemical stability of benzene and the rotational barrier of allyl radical. Further application of the model shows that  $\pi$ -delocalization, per se, is seldom expected to be a driving force in organic molecules containing C, N, and O. In this manner the delocalization problem is unified and shown *not* to be merely a matter of electron count and mode of delocalization.

In recent communications,<sup>2</sup> ab initio computational evidence has been provided that the  $\pi$ -electrons of allyl radical (**1**) and benzene (**2**) prefer to distort to their localized  $\pi$ -components, much like the  $\pi$ -electrons of singlet cyclobutadiene (**3**). These distortive



propensities in **1** and **2** are, however, quenched by the  $\sigma$ -frames that strongly prefer regular geometries with uniform C-C bond lengths. Consequently,  $\pi$ -electronic delocalization in **1** and **2** turns

out to be a byproduct of a geometric constraint and occurs despite the opposite inherent tendency of the  $\pi$ -electrons. This result touches a key question of chemical epistemology: is electronic delocalization a driving force of stability and geometric shape? Our preliminary computational data<sup>2</sup> show that, in organic species,  $\pi$ -delocalization is not such a driving force but rather a byproduct phenomenon.

The credibility of this conclusion must be established now by a two-pronged examination. Firstly, it must be shown that it is not a computational artifact but rather a chemical trend that emerges from the broader context of chemistry. Secondly, the

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